



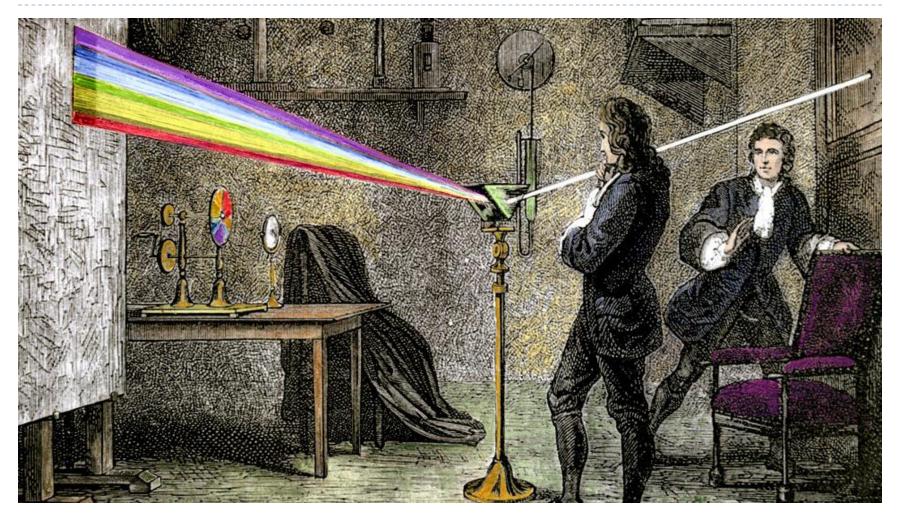
Light and colour

Advanced Graphics

Rafal Mantiuk Computer Laboratory, University of Cambridge

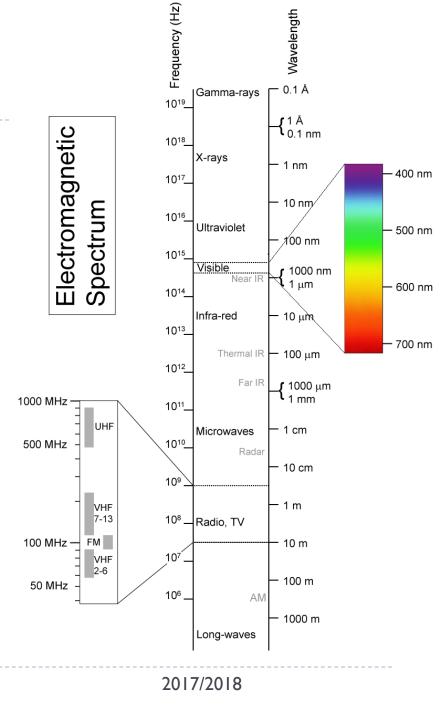
2017/2018

From light to colour spaces



Electromagnetic spectrum

- Visible light
 - Electromagnetic waves of wavelength in the range 380nm to 730nm
 - Earth's atmosphere lets through a lot of light in this wavelength band
 - Higher in energy than thermal infrared, so heat does not interfere with vision



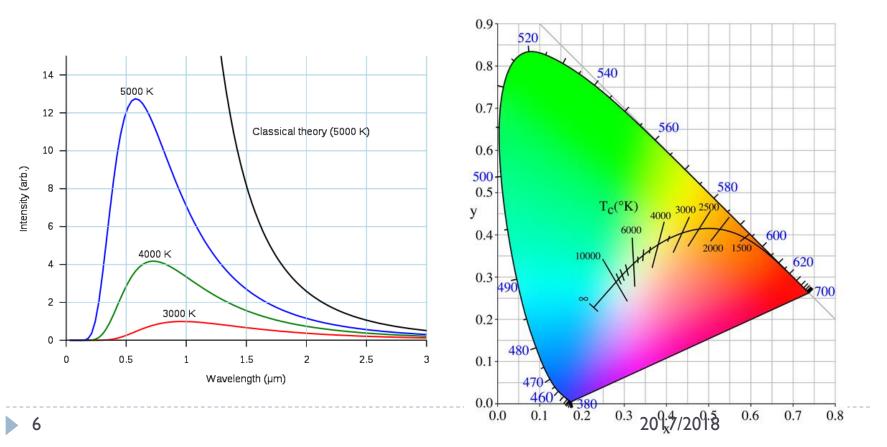
Colour

There is no physical definition of colour – colour is the result of our perception



Black body radiation

- Electromagnetic radiation emitted by a perfect absorber at a given temperature
 - Graphite is a good approximation of a black body



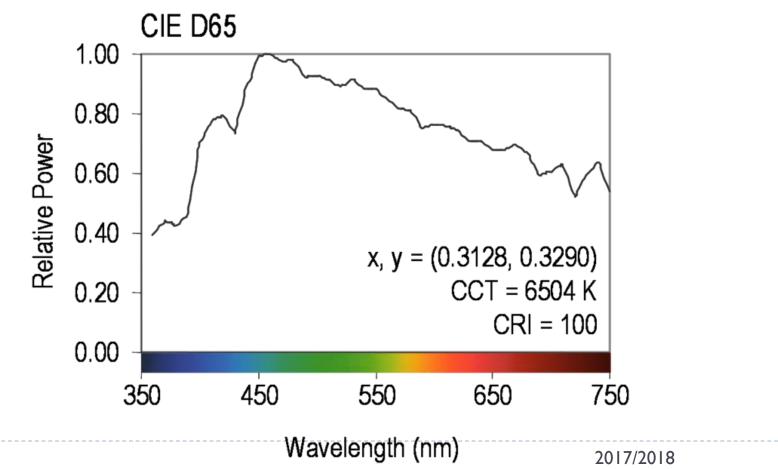
Correlated colour temperature

- The temperature of a black body radiator that produces light most closely matching the particular source
- Examples:
 - Typical north-sky light: 7500 K
 - Typical average daylight: 6500 K
 - Domestic tungsten lamp (100 to 200 W): 2800 K
 - Domestic tungsten lamp (40 to 60 W): 2700 K
 - Sunlight at sunset: 2000 K
- Useful to describe colour of the illumination (source of light)



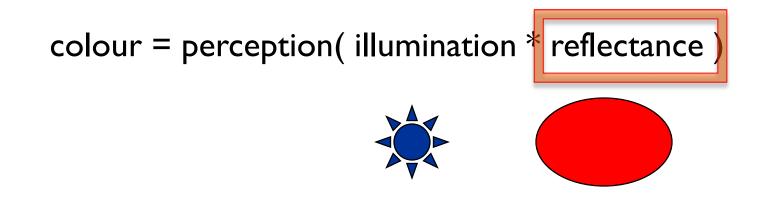


- Mid-day sun in Western Europe / Northern Europe
- Colour temperature approx. 6500 K



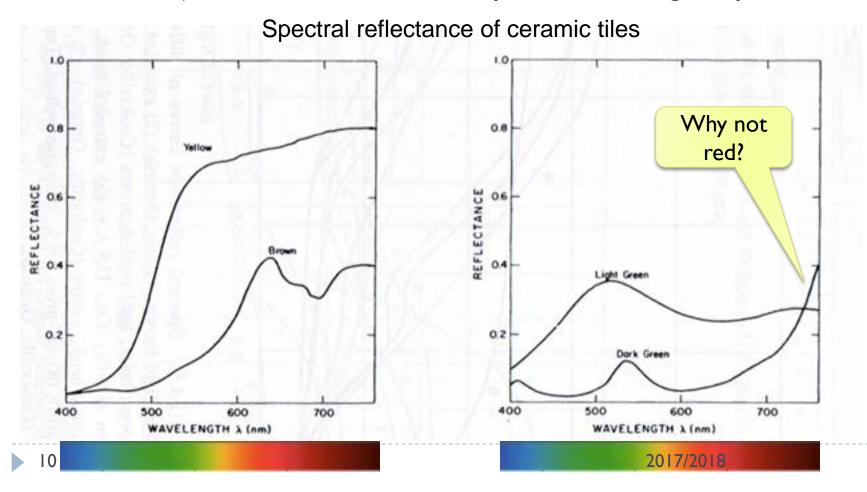
Colour

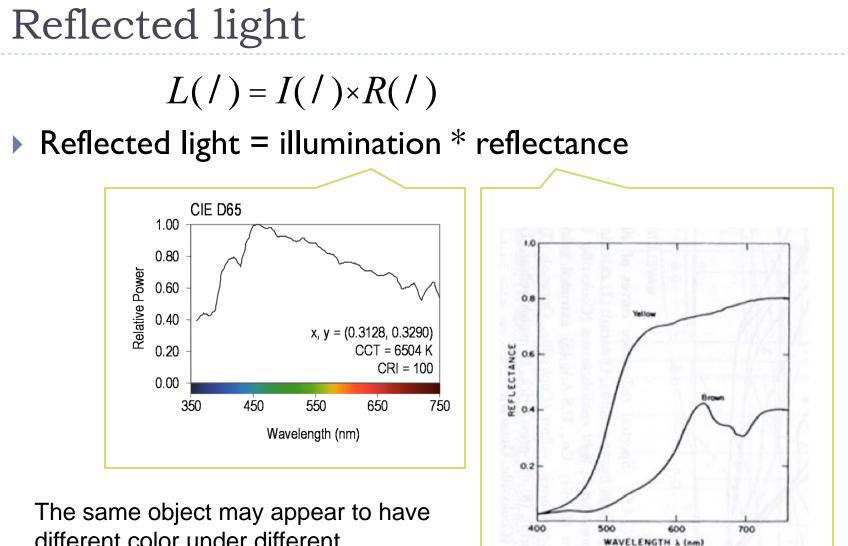
There is no physical definition of colour – colour is the result of our perception



Reflectance

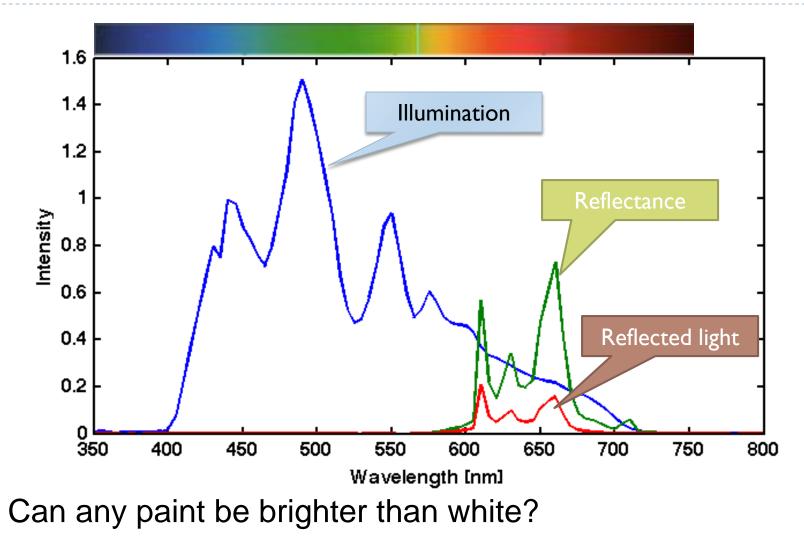
- Most of the light we see is reflected from objects
- These objects absorb a certain part of the light spectrum

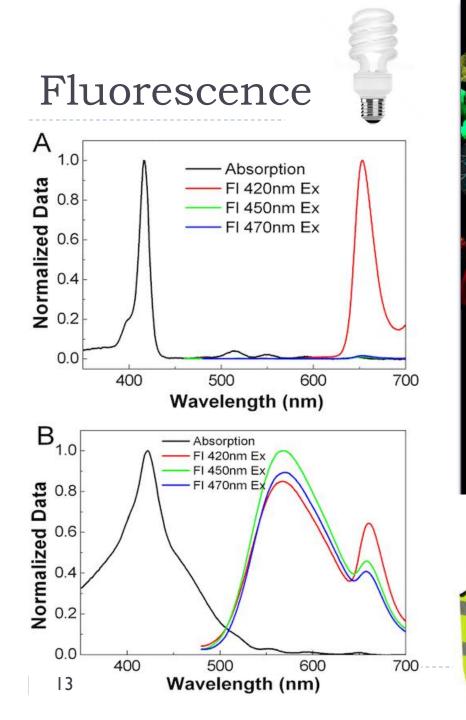




different color under different illumination.

Example

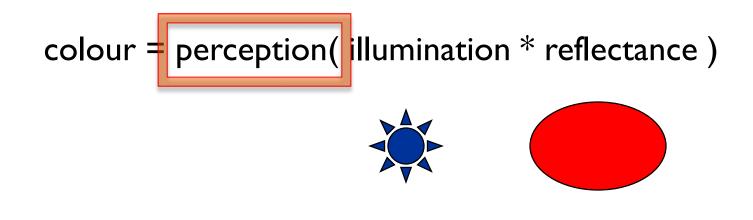






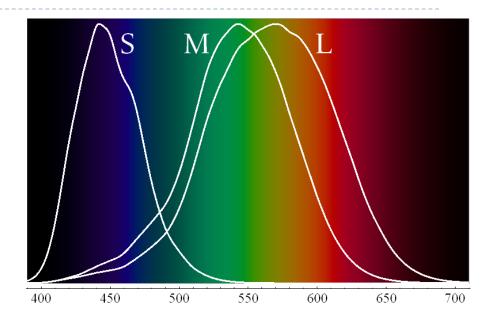
Colour

There is no physical definition of colour – colour is the result of our perception



Colour vision

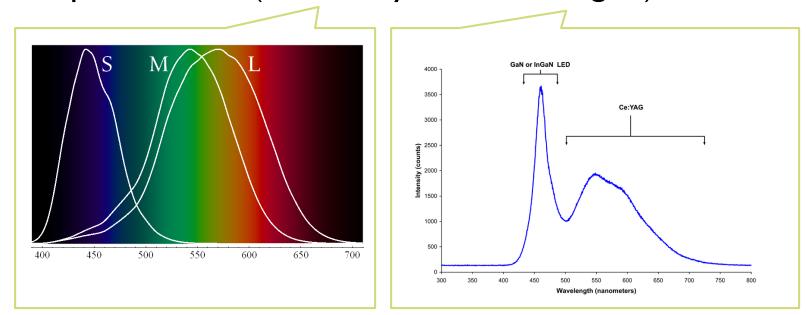
- Cones are the photreceptors responsible for color vision
 - Only daylight, we see no colors when there is not enough light
- Three types of cones
 - S sensitive to short wavelengths
 - M sensitive to medium wavelengths
 - L sensitive to long wavelengths



Sensitivity curves – probability that a photon of that wavelengths will be absorbed by a photoreceptor

Perceived light

cone response = sum(sensitivity * reflected light)



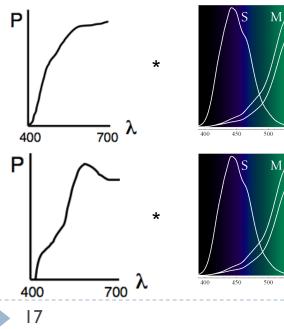
Although there is an infinite number of wavelengths, we have only three photoreceptor types to sense differences between light spectra Formally $R_{S} = \int_{380}^{730} S_{S}(\lambda) \cdot L(\lambda) d\lambda$

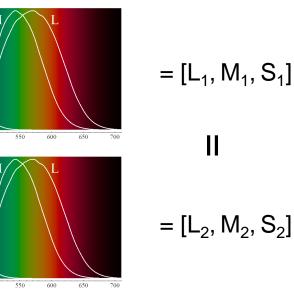
Index S for S-cones

2017/2018

Metamers

- Even if two light spectra are different, they may appear to have the same colour
- The light spectra that appear to have the same colour are called metamers
- Example:

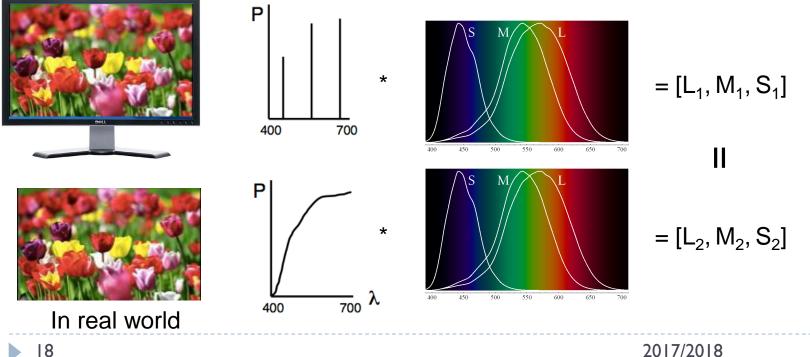




Practical application of metamerism

- Displays do not emit the same light spectra as real-world objects
- Yet, the colours on a display look almost identical

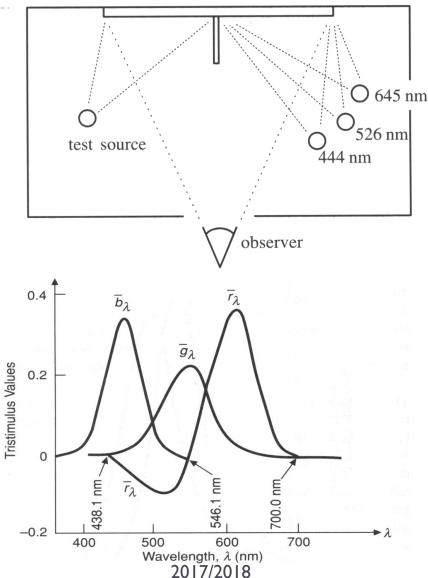
On display



Tristimulus Colour Representation

Observation

- Any colour can be matched using three linear independent reference colours
- May require "negative" contribution to test colour
- Matching curves describe the value for matching monochromatic spectral colours of equal intensity
 - With respect to a certain set of primary colours



Standard Colour Space CIE-XYZ

CIE Experiments [Guild and Wright, 1931]

- Colour matching experiments
- Group ~12 people with ,,normal" colour vision
- 2 degree visual field (fovea only)

CIE 2006 XYZ

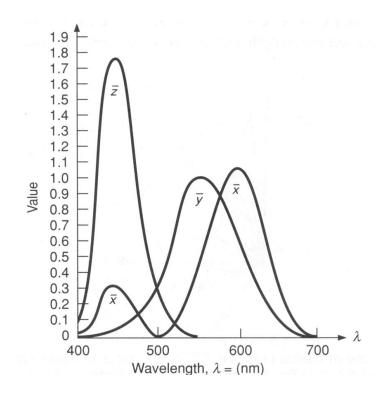
- Derived from LMS color matching functions by Stockman & Sharpe
- S-cone response differs the most from CIE 1931

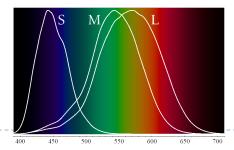
CIE-XYZ Colour Space

- Goals
 - Abstract from concrete primaries used in experiment
 - All matching functions are positive
 - One primary is roughly proportionally to light intensity

Standard Colour Space CIE-XYZ

- Standardized imaginary primaries CIE XYZ (1931)
 - Could match all physically realizable colour stimuli
 - Y is roughly equivalent to luminance
 - Shape similar to luminous efficiency curve
 - Monochromatic spectral colours form a curve in 3D XYZ-space





Cone sensitivity curves can be obtained by a linear transformation of CIE XYZ

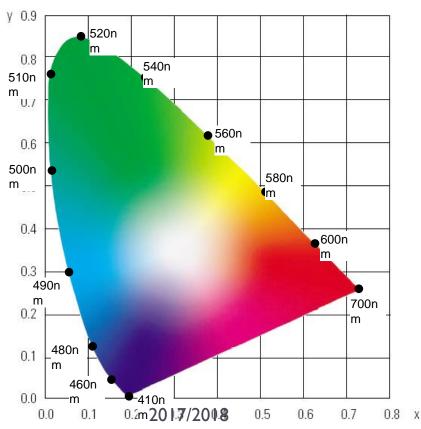
2017/2018

CIE chromaticity diagram

chromaticity values are defined in terms of x, y, z

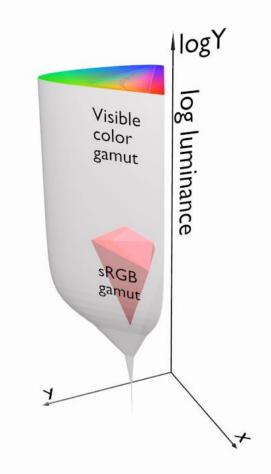
$$x = \frac{X}{X+Y+Z}, \quad y = \frac{Y}{X+Y+Z}, \quad z = \frac{Z}{X+Y+Z}$$
 $\therefore \quad x+y+z = 1$

- ignores luminance
- can be plotted as a 2D function
- pure colours (single wavelength)
 lie along the outer curve
- all other colours are a mix of pure colours and hence lie inside the curve
- points outside the curve do not exist as colours



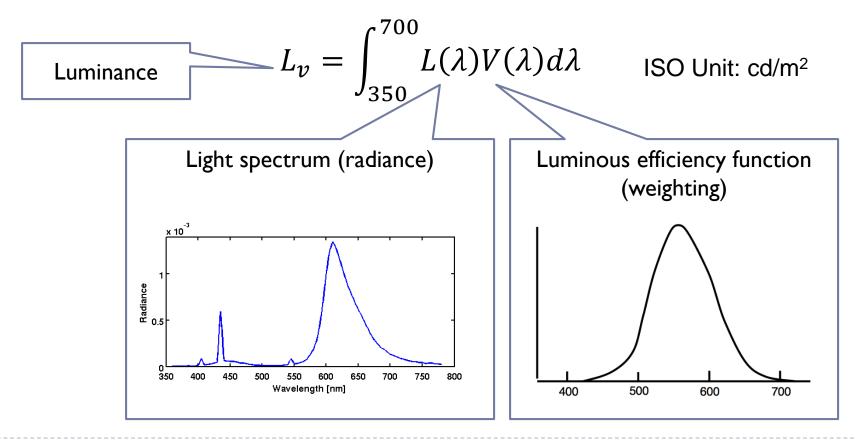
Visible vs. displayable colours

- All physically possible and visible colours form a solid in XYZ space
- Each display device can reproduce a subspace of that space
- A chromacity diagram is a slice taken from a 3D solid in XYZ space
- Colour Gamut the solid in a colour space
 - Usually defined in XYZ to be deviceindependent



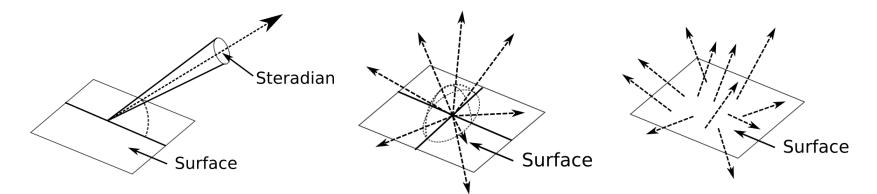
Luminance – photometric quantity

Luminance – perceived brightness of light, adjusted for the sensitivity of the visual system to wavelengths



Photometric units

Quantity	Units	Symbol
Luminance	candela per sq. meter [cd/m²=lm/(sr*m²)]	L _v
Illuminance	$lux [lx = lm/m^2 = cd*sr/m^2]$	E _V
Luminous flux	lumen [lm = cd*sr]	F



Luminance – light emitted from a point on a surface in a particular direction

Illuminance – light emitted from a point on a surface in all directions **Luminous flux** – light emitted from the entire surface in all directions

All these units can measure either incoming or emitted light

Luminous flux - lumens

- Total light emitted
- Useful to measure and compare light sources
 - For example fluorescent and incandescent light bulbs
- But also used for digital projectors



Integrating sphere – to measure all light emitted

Illuminance - lux

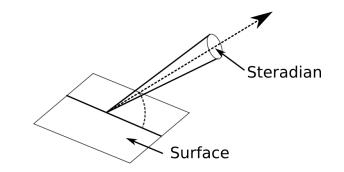
- Measures light coming (or emitted) from all directions
- Useful to measure lighting conditions
 - Whether street lighting is bright enough, etc.



Illuminance meter

Luminance – candela per square meter

- Light emitted (or incomming) from a point in a particular direction
- Luminance is the same regardless of the distance to the emitter
- The light sensed by our eyes is relative to luminance

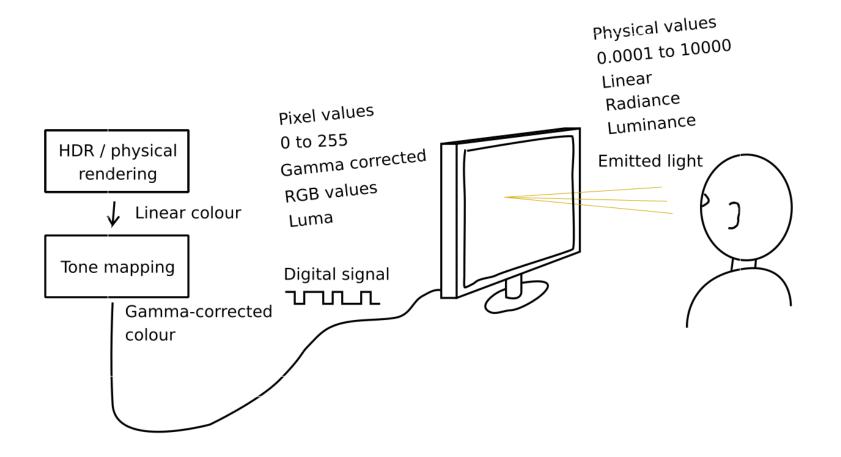


Radiometric vs. Photometric units

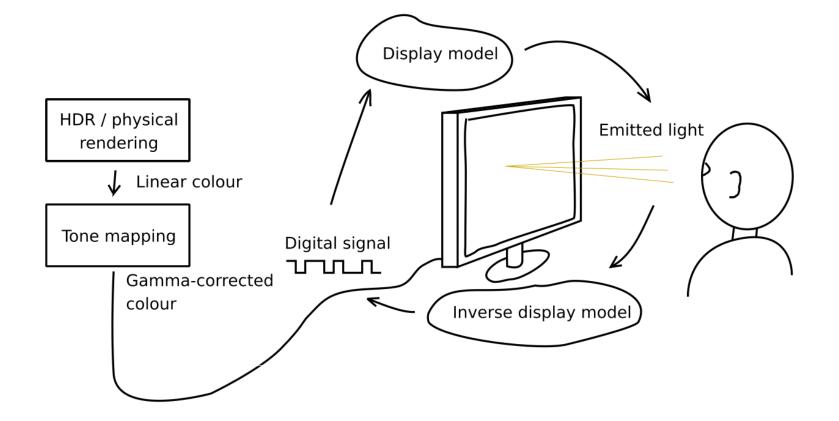
Photometry	Radiometry
Luminance [cd m ⁻²]	Radiance [W sr ⁻¹ m ⁻²]
Illuminance [$Ix = Im m^{-2} = cd sr m^{-2}$]	Irradiance / Exitance / Radiosity [W m ⁻²]
Luminous flux [lm = cd sr]	Radiant flux [W]

- Radiometric units integrate light over all wavelengths (visible and invisible)
- Spectral radiance / irradiance / radiant flux describe light for a single wavelength
- But, in computer graphics radiometric units are often assumed to capture a quantity integrated over a spectral basis function (e.g. red, green, blue)
- In color science, the product of radiance with a colour matching function is called trichromatic colour value

Linear vs. gamma-corrected values



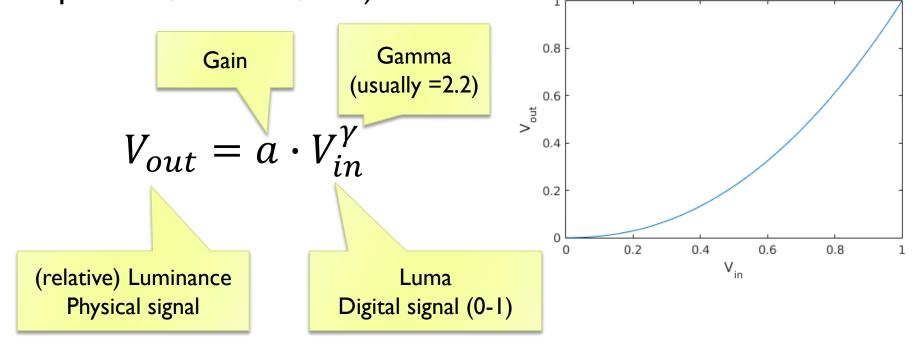
Linear vs. gamma-corrected values



D

Basic display model: gamma correction

 Gamma correction is used to encode luminance or tristimulus color values (RGB) in imaging systems (displays, printers, cameras, etc.)



For color images: $R = a \cdot (R')^{\gamma}$ and the same for green and blue

Why is gamma needed?

 Linear encoding V_S =
 0.0
 0.1
 0.2
 0.3
 0.4
 0.6
 0.7
 0.8
 0.9
 1.0

 Linear intensity
 I =
 0.0
 0.1
 0.3
 0.4
 0.5
 0.6
 0.7
 0.8
 0.9
 1.0

<- Pixel value (luma)

<- Luminance

- "Gamma corrected" pixel values give a scale of brightness levels that is more perceptually uniform
- At least 12 bits (instead of 8) would be needed to encode each color channel without gamma correction
- And accidentally it was also the response of the CRT gun

Luma – gray-scale pixel value

• Luma - pixel brightness in gamma corrected units L' = 0.2126R' + 0.7152G' + 0.0722B'

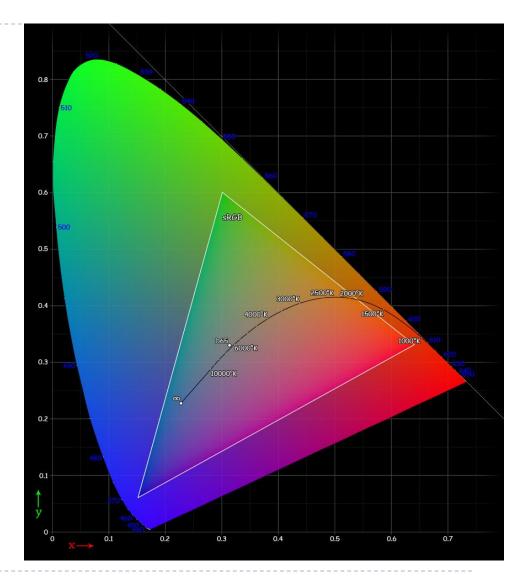
- ▶ *R*′, *G*′ and *B*′ are gamma corrected colour values
- Prime symbol denotes "gamma corrected"
- Used in image/video coding
- Note that relative luminance if often approximated with L = 0.2126R + 0.7152G + 0.0722B $= 0.2126(R')^{\gamma} + 0.7152(G')^{\gamma} + 0.0722(B')^{\gamma}$
- \triangleright R, G, and B are linear colour values
- Luma and luminace are different quantities despite similar formulas

sRGB color space

- "RGB" color space is not a standard. Colors may differ depending on the choice of the primaries
- "sRGB" is a standard color space, which most displays try to mimic (standard for HDTV)

Chromaticity	Red	Green	Blue	White point
x	0.6400	0.3000	0.1500	0.3127
У	0.3300	0.6000	0.0600	0.3290
z	0.0300	0.1000	0.7900	0.3583

 The chromacities above are also known as Rec. 709



sRGB color space

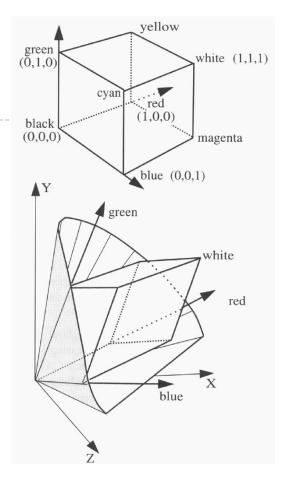
Two step XYZ – sRGB transformation:

Step I: Linear color transform

R_{linear}		3.2406	-1.5372	-0.4986	$\begin{bmatrix} X \end{bmatrix}$
G_{linear}	=	-0.9689	1.8758	0.0415	Y
B_{linear}		0.0557	-1.5372 1.8758 -0.2040	1.0570	$\lfloor Z \rfloor$

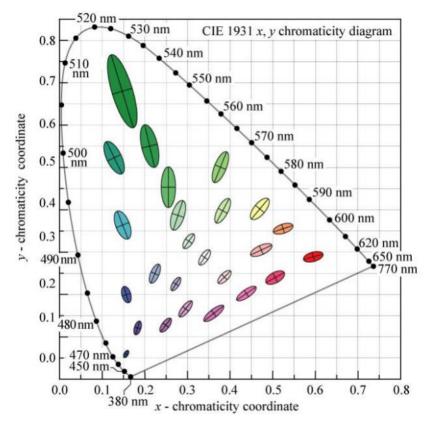
$$C_{\text{srgb}} = \begin{cases} 12.92C_{\text{linear}}, & C_{\text{linear}} \le 0.0031308\\ (1+a)C_{\text{linear}}^{1/2.4} - a, & C_{\text{linear}} > 0.0031308 \end{cases}$$

a = 0.055

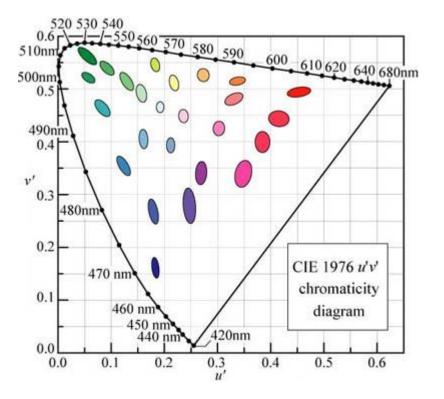


Perceptually uniformity

MacAdam ellipse - visually indistinguishable colours







In CIE u'v' chromatic coordinates

CIE L^{*}u^{*}v^{*} and u'v'

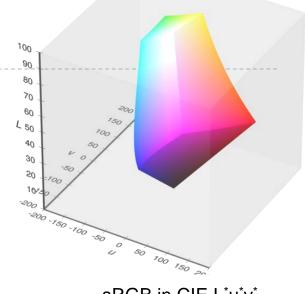
Approximately perceptually uniform
u'v' chromacity

$$u' = \frac{4X}{X + 15Y + 3Z} = \frac{4x}{-2x + 12y + 3}$$
$$v' = \frac{9Y}{X + 15Y + 3Z} = \frac{9y}{-2x + 12y + 3}$$
$$\bullet \text{ CIE LUV}$$

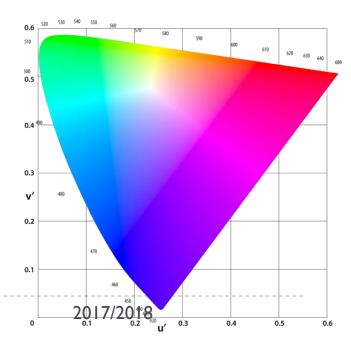
Lightness
$$L^* = \begin{cases} \left(\frac{29}{3}\right)^3 Y/Y_n, & Y/Y_n \le \left(\frac{6}{29}\right)^3 \\ 116(Y/Y_n)^{1/3} - 16, & Y/Y_n > \left(\frac{6}{29}\right)^3 \end{cases}$$
Chromacity
coordinates $u^* = 13L^* \cdot (u' - u'_n) \\ v^* = 13L^* \cdot (v' - v'_n) \end{cases}$ Colours less
distinguishable
when dark

• Hue and chroma

$$egin{aligned} C^*_{uv} &= \sqrt{(u^*)^2 + (v^*)^2} \ h_{uv} &= \mathrm{atan2}(v^*, u^*), \end{aligned}$$



sRGB in CIE L^{*}u^{*}v^{*}



CIE L*a*b* colour space

 Another approximately perceptually uniform colour space

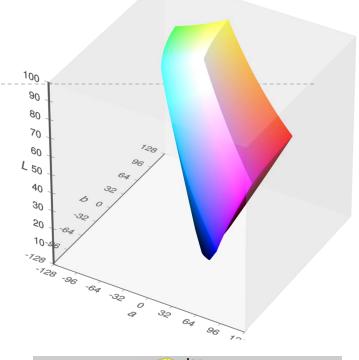
$$\begin{split} L^{\star} &= 116f\left(\frac{Y}{Y_{\rm n}}\right) - 16 \\ a^{\star} &= 500\left(f\left(\frac{X}{X_{\rm n}}\right) - f\left(\frac{Y}{Y_{\rm n}}\right)\right) \\ b^{\star} &= 200\left(f\left(\frac{Y}{Y_{\rm n}}\right) - f\left(\frac{Z}{Z_{\rm n}}\right)\right) \\ b^{\star} &= 100,000, \\ I_{\rm n} &= 108.883 \end{split}$$

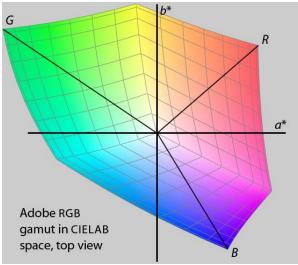
$$f(t) &= \begin{cases} \sqrt[3]{t} & \text{if } t > \delta^3 \\ \frac{t}{3\delta^2} + \frac{4}{29} & \text{otherwise} \end{cases}$$

$$\delta &= \frac{6}{29} \end{split}$$

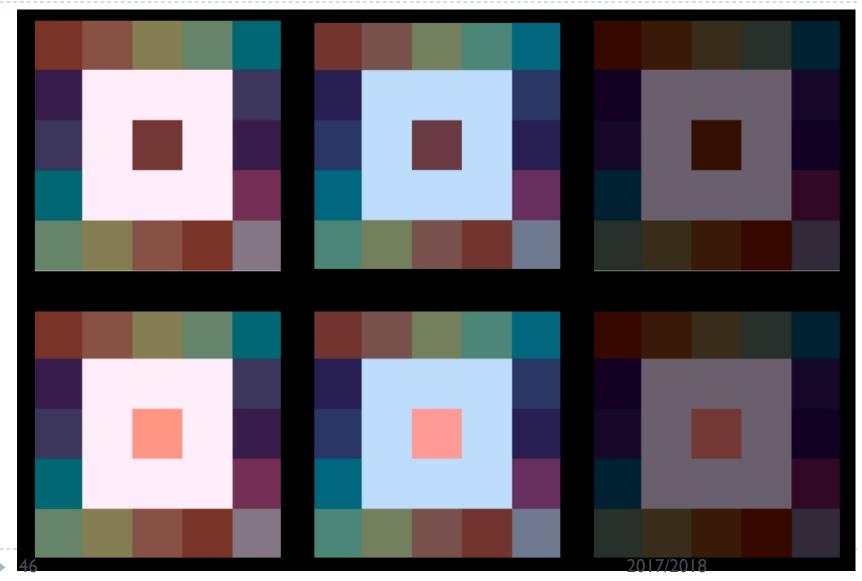
Chroma and hue

$$C^{\star} = \sqrt{{a^{\star}}^2 + {b^{\star}}^2}, \qquad h^\circ = rctaniggl(rac{b^{\star}}{a^{\star}}iggr)$$





Colour constancy



from: http://www.uv.es/vista/vistavalencia/color/color.html

Chromatic adaptation = colour constancy

- Visual system "estimates" the colour of the illuminant
 - and then attempts to discount it
- This works well if the scene fills the entire field of view
 - But is less effective for images
 - E.g. image on the computer monitor or developed print
- Therefore photographs require white balance
 - To discount the illuminant that is not discounted by the visual system



from Wikipedia



2017/2018

White point

Displays are expected to have the white point D65

- This corresponds to the color temperature of 6500K
- But most displays do not strictly adhere to this specification
- It is often possible to adjust the white point of a display
- Digital cameras need to discount illuminant
 - They estimate the color of white and make it D65 so that it looks white on displays
 - This is called white balance



From: http://en.wikipedia.org/wiki/File:Incand-3500-5500-color-temp-comparison.png

References

- Well written textbook
 - Fairchild, M. D. (2005). Color Appearance Models (second.). John Wiley & Sons.
- More detailed introduction to light and colour phenomena
 - Erik Reinhard, Erum Arif Khan, Ahmet Oguz Akyuz, G. J. (2008). Color Imaging: Fundamentals and Applications. CRC Press.