

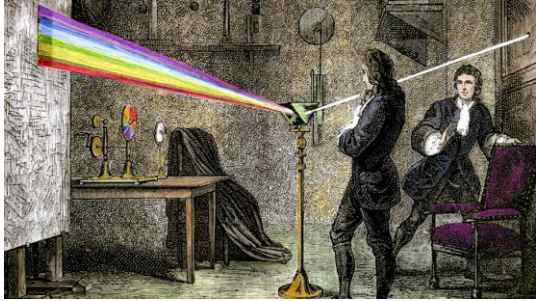
Light and colour

Advanced Graphics

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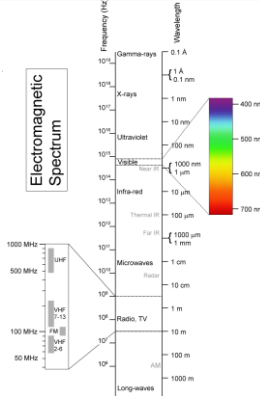
From light to colour spaces



2

Electromagnetic spectrum

- ▶ **Visible light**
 - ▶ Electromagnetic waves of wavelength in the range 380nm to 730nm
 - ▶ Earth's atmosphere lets through a lot of light in this wavelength band
 - ▶ Higher in energy than thermal infrared, so heat does not interfere with vision




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Colour

- ▶ There is no physical definition of colour – colour is the result of our perception

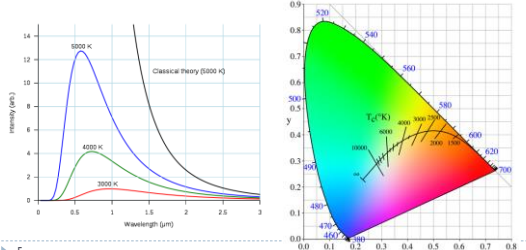
colour = perception (illumination * reflectance)



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Black body radiation


- ▶ Electromagnetic radiation emitted by a perfect absorber at a given temperature
- ▶ Graphite is a good approximation of a black body



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Correlated colour temperature

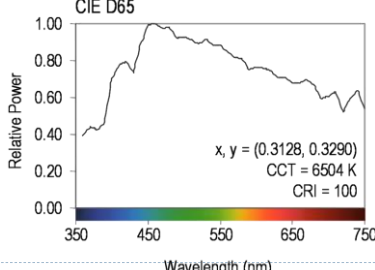
- ▶ The temperature of a black body radiator that produces light most closely matching the particular source
- ▶ Examples:
 - ▶ Typical north-sky light: 7500 K
 - ▶ Typical average daylight: 6500 K
 - ▶ Domestic tungsten lamp (100 to 200 W): 2800 K
 - ▶ Domestic tungsten lamp (40 to 60 W): 2700 K
 - ▶ Sunlight at sunset: 2000 K
- ▶ Useful to describe colour of the **illumination** (source of light)



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Standard illuminant D65

- ▶ Mid-day sun in Western Europe / Northern Europe
- ▶ Colour temperature approx. 6500 K




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Colour

- ▶ There is no physical definition of colour – colour is the result of our perception

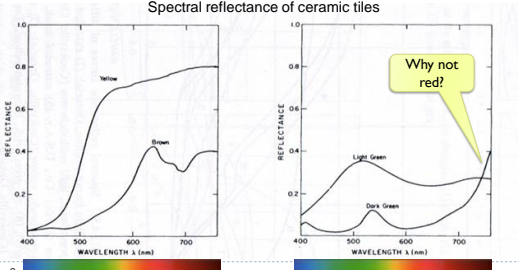
colour = perception(illumination * reflectance)



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Reflectance

- ▶ Most of the light we see is reflected from objects
- ▶ These objects absorb a certain part of the light spectrum

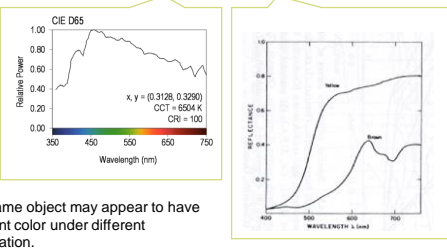


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Reflected light

$$L(\lambda) = I(\lambda) \times R(\lambda)$$

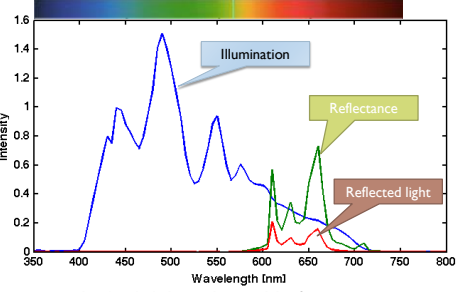
- ▶ Reflected light = illumination * reflectance



The same object may appear to have different color under different illumination.

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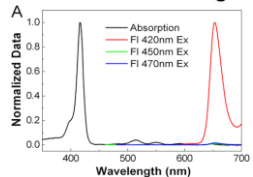
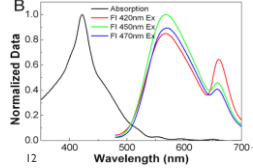
Example



Can any paint be brighter than white?

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Fluorescence

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Colour

- ▶ There is no physical definition of colour – colour is the result of our perception

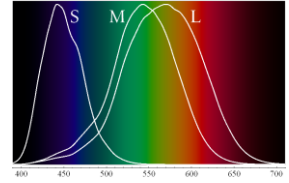
colour = perception(illumination * reflectance)



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Colour vision

- ▶ Cones are the photoreceptors responsible for color vision
 - ▶ Only daylight, we see no colors when there is not enough light
- ▶ Three types of cones
 - ▶ S – sensitive to short wavelengths
 - ▶ M – sensitive to medium wavelengths
 - ▶ L – sensitive to long wavelengths

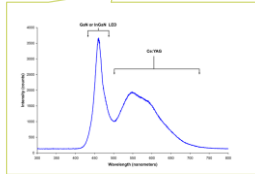
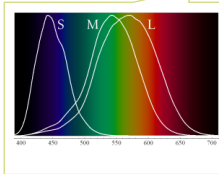


Sensitivity curves – probability that a photon of that wavelengths will be absorbed by a photoreceptor

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Perceived light

- ▶ cone response = sum(sensitivity * reflected light)



Although there is an infinite number of wavelengths, we have only three photoreceptor types to sense differences between light spectra

Formally

$$R_s = \int_{380}^{730} S_s(\lambda) \cdot L(\lambda) d\lambda$$

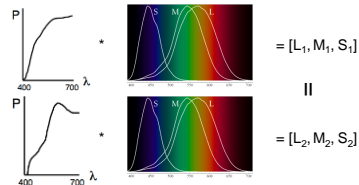
Index S for S-cones

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Metamers

- ▶ Even if two light spectra are different, they may appear to have the same colour
- ▶ The light spectra that appear to have the same colour are called **metamers**

- ▶ Example:

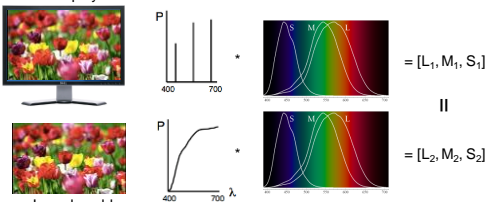


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Practical application of metamerism

- ▶ Displays do not emit the same light spectra as real-world objects
- ▶ Yet, the colours on a display look almost identical

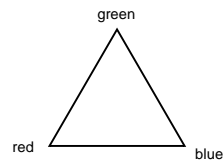
On display



In real world

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Tristimulus Colour Representation

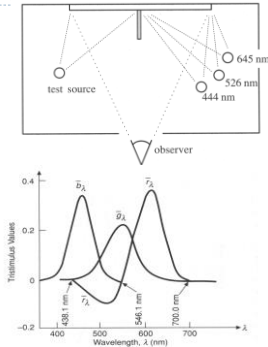


- ▶ Interpolation of primaries yields triangle of colours
- ▶ Making use of the three cones and their weighting functions

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Tristimulus Colour Representation

- ▶ **Observation**
 - ▶ Any colour can be matched using three linear independent reference colours
 - ▶ May require "negative" contribution to test colour
 - ▶ Matching curves describe the value for matching monochromatic spectral colours of equal intensity
 - ▶ With respect to a certain set of primary colours



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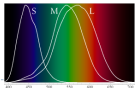
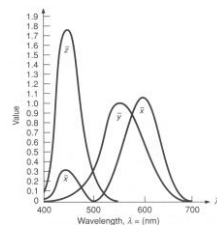
Standard Colour Space CIE-XYZ

- ▶ **CIE Experiments [Guild and Wright, 1931]**
 - ▶ Colour matching experiments
 - ▶ Group ~12 people with „normal“ colour vision (from London area)
 - ▶ 2 degree visual field (fovea only)
 - ▶ Other Experiment in 1964
 - ▶ 10 degree visual field, ~50 people (with foreigners)
 - ▶ More appropriate for larger field of view but rarely used
- ▶ **CIE-XYZ Colour Space**
 - ▶ Goals
 - ▶ Abstract from concrete primaries used in experiment
 - ▶ All matching functions are positive
 - ▶ One primary is roughly proportionally to light intensity

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Standard Colour Space CIE-XYZ

- ▶ **Standardized imaginary primaries CIE XYZ (1931)**
 - ▶ Could match all physically realizable colour stimuli
 - ▶ Y is roughly equivalent to luminance
 - ▶ Shape similar to luminous efficiency curve
 - ▶ Monochromatic spectral colours form a curve in 3D XYZ-space

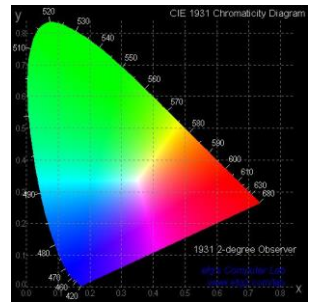


Cone sensitivity curves can be obtained by a linear transformation of CIE XYZ

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CIE Chromacity diagram

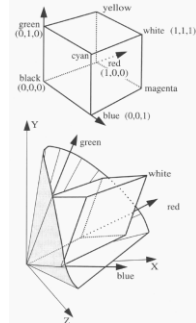
- **Normalization:**
 - Concentrate on colour, not light intensity
 - Relative colour coordinates
 - $x = \frac{X}{X+Y+Z}$ etc
- Chromaticity diagram: 2D-Plot over x and y
- Points in diagram are called „colour locations“
- White point: ~ (0.3, 0.3)
 - Device dependent
 - Adaptation of the eye



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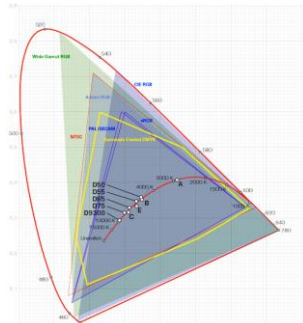
Monitor Color Gamut

- ▶ **CIE XYZ gamut**
 - ▶ Device-independent
- ▶ **Device color gamut**
 - ▶ Cube inside CIE color space with additive color blending



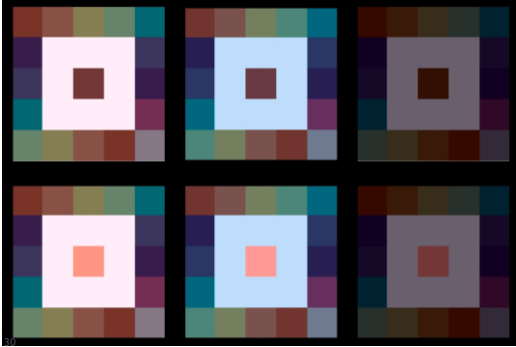
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Different Color Gamuts



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Colour constancy



from: <http://www.uv.es/vista/vistavalencia/color/color.html>

Chromatic adaptation = colour constancy

- ▶ Visual system “estimates” the colour of the illuminant
 - ▶ and then attempts to discount it
- ▶ This works well if the scene fills the entire field of view
 - ▶ But is less effective for images
 - ▶ E.g. image on the computer monitor or developed print
- ▶ Therefore photographs require **white balance**
 - ▶ To discount the illuminant that is not discounted by the visual system



from Wikipedia

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White point

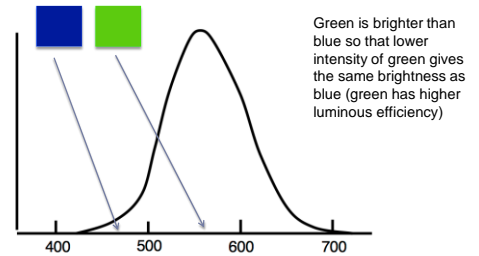
- ▶ Displays are expected to have the white point D65
 - ▶ This corresponds to the color temperature of 6500K
 - ▶ But most displays do not strictly adhere to this specification
 - ▶ It is often possible to adjust the white point of a display
- ▶ Digital cameras need to discount illuminant
 - ▶ They estimate the color of white and make it D65 so that it looks white on displays
 - ▶ This is called white balance



From: <http://en.wikipedia.org/wiki/File:Incand-3500-5500-color-temp-comparison.png>

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Luminous efficiency function



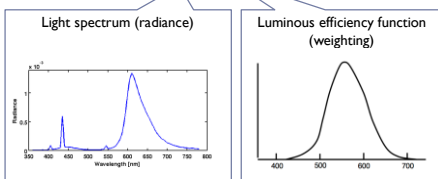
To match the brightness of colors produced by the light of different wavelength

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Photometric units

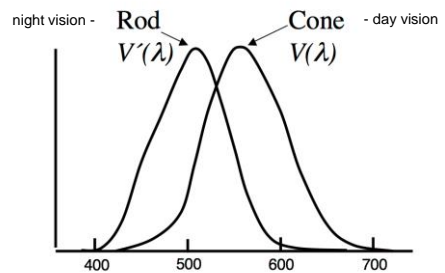
- ▶ Luminance – perceived brightness of light, adjusted for the sensitivity of the visual system to wavelengths

$$L_v = \int_0^\infty L(\lambda) \cdot V(\lambda) d\lambda \quad \text{ISO Unit: cd/m}^2$$



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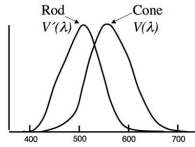
Rod and cone luminous efficiency functions



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Purkinje shift (effect)

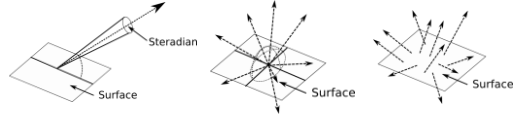
- ▶ A shift in spectral sensitivity associated with the transition of cone to rod vision
- ▶ Blue appears brighter and red appears darker in twilight
- ▶ And the reverse is observed in daylight



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Photometric units

Quantity	Units	Symbol
Luminance	candela per sq. meter [$\text{cd}/\text{m}^2 = \text{lm}/(\text{sr} \cdot \text{m}^2)$]	L_v
Illuminance	lux [$\text{lx} = \text{lm}/\text{m}^2 = \text{cd} \cdot \text{sr}/\text{m}^2$]	E_v
Luminous flux	lumen [$\text{lm} = \text{cd} \cdot \text{sr}$]	F



Luminance – light emitted from a point on a surface in a particular direction

Illuminance – light emitted from a point on a surface in all directions

Luminous flux – light emitted from the entire surface in all directions

All these units can measure either incoming or emitted light

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Luminous flux - lumens

- ▶ Total light emitted
- ▶ Useful to measure and compare light sources
 - ▶ For example fluorescent and incandescent light bulbs
- ▶ But also used for digital projectors



Integrating sphere – to measure all light emitted

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Illuminance - lux

- ▶ Measures light coming (or emitted) from all directions
- ▶ Useful to measure lighting conditions
 - ▶ Whether street lighting is bright enough, etc.

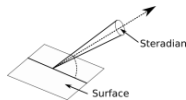


Illuminance meter

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Luminance – candela per square meter

- ▶ Light emitted (or incoming) from a point in a particular direction
- ▶ Luminance is the same regardless of the distance to the emitter
- ▶ The light sensed by our eyes is relative to luminance



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Radiometric vs. Photometric units

Photometry	Radiometry
Luminance [cd/m^2]	Radiance [$\text{W}/\text{sr} \cdot \text{m}^2$]
Illuminance [$\text{lx} = \text{lm}/\text{m}^2 = \text{cd} \cdot \text{sr}/\text{m}^2$]	Irradiance / Exitance / Radiosity [W/m^2]
Luminous flux [$\text{lm} = \text{cd} \cdot \text{sr}$]	Radiant flux [W]

- ▶ Radiometric units integrate light over **all wavelengths** (visible and invisible)
- ▶ **Spectral radiance / irradiance / radiant flux** describe light for a single wavelength
- ▶ But, in computer graphics radiometric units are often assumed to capture a quantity integrated over a spectral basis function (e.g. red, green, blue)
- ▶ In color science, the product of radiance with a colour matching function is called **trichromatic colour value**

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Gamma correction

▶ Gamma correction is used to encode luminance or tristimulus color values (RGB) in imaging systems (displays, printers, cameras, etc.)

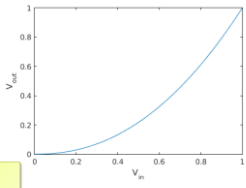
Gain

Gamma (usually =2.2)

$$V_{out} = a \times V_{in}^g$$

(relative) Luminance

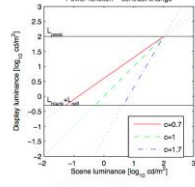
Luma




For color images: $R = a \times (R_c)^g$ and the same for green and blue

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Gamma






(b) $c=0.7$

Lower gamma



(c) $c=1$

Original gamma

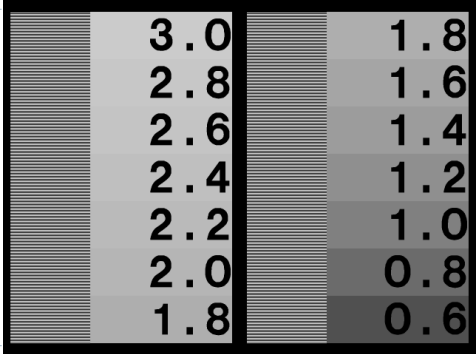


(d) $c=1.7$

Higher gamma

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Gamma Testing Chart



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Why is gamma needed?

Linear encoding $V_S = 0.0 \ 0.1 \ 0.2 \ 0.3 \ 0.4 \ 0.5 \ 0.6 \ 0.7 \ 0.8 \ 0.9 \ 1.0$ <- Pixel value (luma)

Linear intensity $I = 0.0 \ 0.1 \ 0.2 \ 0.3 \ 0.4 \ 0.5 \ 0.6 \ 0.7 \ 0.8 \ 0.9 \ 1.0$ <- Luminance

- ▶ “Gamma corrected” pixel values give a scale of brightness levels that is more perceptually uniform
- ▶ At least 12 bits (instead of 8) would be needed to encode each color channel without gamma correction
- ▶ And accidentally it was also the response of the CRT gun

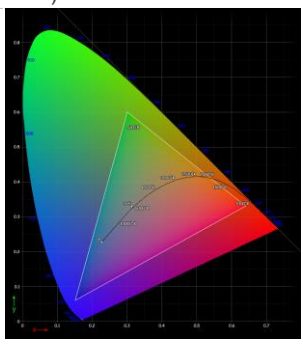
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sRGB color space (LDR)

- ▶ “RGB” color space is not a standard. Colors may differ depending on the choice of the primaries
- ▶ “sRGB” is a standard color space, which most displays try to mimic (standard for HDTV)

Chromaticity	Red	Green	Blue	White point
x	0.6400	0.3000	0.1500	0.3127
y	0.3300	0.6000	0.0600	0.3290
z	0.0300	0.1000	0.7900	0.3583

▶ The chromaticities above are also known as Rec. 709



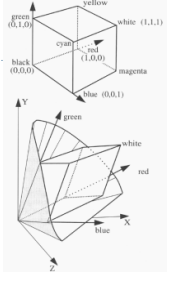
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sRGB color space

- ▶ Two step XYZ – sRGB transformation:
 - ▶ Step 1: Linear color transform

$$\begin{bmatrix} R_{linear} \\ G_{linear} \\ B_{linear} \end{bmatrix} = \begin{bmatrix} 3.2406 & -1.5372 & -0.4986 \\ -0.9689 & 1.8758 & 0.0415 \\ 0.0557 & -0.2040 & 1.0570 \end{bmatrix} \begin{bmatrix} X \\ Y \\ Z \end{bmatrix}$$

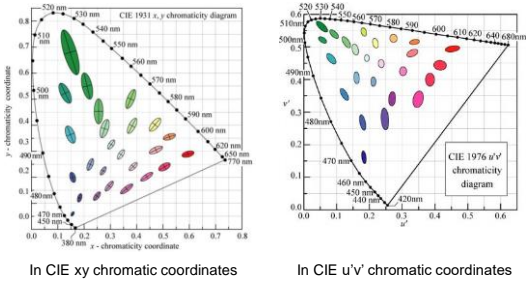
- ▶ Step 2: Non-linearity

$$C_{srgb} = \begin{cases} 12.92 C_{linear}, & C_{linear} \leq 0.0031308 \\ (1 + a) C_{linear}^{1/2.4} - a, & C_{linear} > 0.0031308 \end{cases}$$


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Perceptually uniformity

- ▶ MacAdam ellipse - visually indistinguishable colours



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CIE L*u*v* and u'v'

- ▶ Approximately perceptually uniform
- ▶ u'v' chromaticity

$$u' = \frac{4X}{X + 15Y + 3Z} = \frac{4x}{-2x + 12y + 3}$$

$$v' = \frac{9Y}{X + 15Y + 3Z} = \frac{9y}{-2x + 12y + 3}$$

- ▶ CIE LUV

Lightness $L^* = \begin{cases} \left(\frac{29}{3}\right)^3 Y/Y_n, & Y/Y_n \leq \left(\frac{6}{29}\right)^3 \\ 116(Y/Y_n)^{1/3} - 16, & Y/Y_n > \left(\frac{6}{29}\right)^3 \end{cases}$

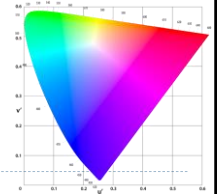
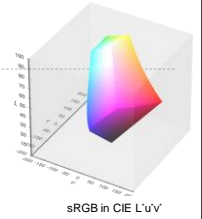
Chromaticity coordinates $u^* = 13L^* \cdot (u' - u'_n)$
 $v^* = 13L^* \cdot (v' - v'_n)$

Colours less distinguishable when dark

- ▶ Hue and chroma

$$C_{uv}^* = \sqrt{(u^*)^2 + (v^*)^2}$$

$$h_{uv} = \text{atan2}(v^*, u^*),$$



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CIE L*a*b* colour space

- ▶ Another approximately perceptually uniform colour space

$$L^* = 116f\left(\frac{Y}{Y_n}\right) - 16$$

$$a^* = 500\left(f\left(\frac{X}{X_n}\right) - f\left(\frac{Y}{Y_n}\right)\right)$$

$$b^* = 200\left(f\left(\frac{Y}{Y_n}\right) - f\left(\frac{Z}{Z_n}\right)\right)$$

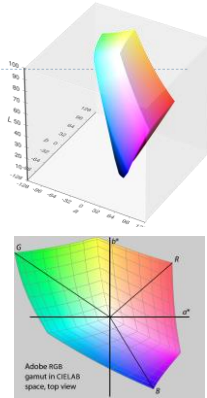
Trichromatic values of the white point, e.g.
 $X_n = 95.047,$
 $Y_n = 100.000,$
 $Z_n = 108.883$

$$f(t) = \begin{cases} \sqrt[3]{t} & \text{if } t > \delta^3 \\ \frac{t}{3\delta} + \frac{4}{29} & \text{otherwise} \end{cases}$$

$$\delta = \frac{6}{29}$$

- ▶ Chroma and hue

$$C^* = \sqrt{a^{*2} + b^{*2}}, \quad h^{\circ} = \arctan\left(\frac{b^*}{a^*}\right)$$



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References

- ▶ Well written textbook
 - ▶ Fairchild, M. D. (2005). *Color Appearance Models* (second.). John Wiley & Sons.
- ▶ More detailed introduction to light and colour phenomena
 - ▶ Erik Reinhard, Erum Arif Khan, Ahmet Oguz Akyuz, G. J. (2008). *Color Imaging: Fundamentals and Applications*. CRC Press.

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